from the expected distribution $\omega(\varepsilon_1, \varepsilon_2, \varepsilon_3)$ defined in Section 11.3.2.3 and removed from the final background value.

11.3.3.3. *Standard profiles*

Reflection profiles are represented on the Ewald sphere within a domain D_0 comprising $2n_1+1, 2n_2+1, 2n_3+1$ equidistant grid-points along $\mathbf{e}_1, \mathbf{e}_2, \mathbf{e}_3$, respectively. The sampling distances between adjacent grid points are then $\Delta_1 = \delta_D/(2n_1+1)$, $\Delta_2 = \delta_D/(2n_2+1), \Delta_3 = \delta_M/(2n_3+1)$. Thus, grid coordinate ν_3 $(\nu_3 = -n_3, \ldots, n_3)$ covers the set of rotation angles

$$\Gamma_{\nu_3} = \{\varphi' | (\nu_3 - 1/2)\Delta_3 \le (\varphi' - \varphi)\cdot\zeta \le (\nu_3 + 1/2)\Delta_3\}.$$

Contributions to the spot intensity come from one or several adjacent data images $(j = j_1, \ldots, j_2)$, each covering the set of rotation angles

$$\Gamma_j = \{\varphi' | \varphi_0 + (j-1)\Delta_\varphi \le \varphi' \le \varphi_0 + j\Delta_\varphi\}.$$

Assuming Gaussian profiles along $\mathbf{e}_3$ for all reflections (see Section 11.3.2.3), the fraction of counts (after subtraction of the background) contributed by data frame j to grid coordinate ν_3 is

$$f_{\nu_3 j} \approx \int\limits_{\Gamma_j \cap \Gamma_{\nu_3}} \exp[-(\varphi' - \varphi)^2/2\sigma^2]\,\mathrm{d}\varphi' \times \left\{ \int\limits_{\Gamma_j} \exp[-(\varphi' - \varphi)^2/2\sigma^2]\,\mathrm{d}\varphi' \right\}^{-1},$$

where $\sigma = \sigma_M/|\zeta|$. The integrals can be expressed in terms of the error function, for which efficient numerical approximations are available (Abramowitz & Stegun, 1972). Finally, each pixel on data image j belonging to the reflection is subdivided into 5×5 areas of equal size, and $f_{\nu_3 j}/25$ of the pixel signal is added to the profile value at grid coordinates ν_1, ν_2, ν_3 corresponding to each subdivision.

This complicated procedure leads to more uniform intensity profiles for all reflections than using their untransformed shape. This simplifies the task of modelling the expected intensity distribution needed for integration by profile fitting. As implemented in *XDS*, reference profiles are learnt every 5° of crystal rotation at nine positions on the detector, each covering an equal area of the detector face. In the learning phase, profile boxes of the strong reflections are normalized and added to their nearest reference profile boxes. The contributions are weighted according to the distance from the location of the reference profile. Each grid point within the average profile boxes is classified as signal if it is above 2% of the peak maximum. Finally, each profile is scaled such that the sum of its signal pixels normalizes to one. The analytic expression $\omega(\varepsilon_1, \varepsilon_2, \varepsilon_3)$ defined in Section 11.3.2.3 for the expected intensity distribution is only a rough initial approximation which is now replaced by the empirical reference profiles.

11.3.3.4. *Intensity estimation*

If an expected intensity distribution $\{p_i | i \in D_0\}$ of the observed profile is given in a domain D_0, the reflection intensity I can be estimated as

$$I = \sum_{i\in D}(c_i - b_i)p_i/v_i \bigg/ \sum_{i\in D} p_i^2/v_i,$$

which minimizes the function

$$\psi(I) = \sum_{i\in D}(c_i - I\cdot p_i - b_i)^2/v_i, \qquad \sum_{i\in D_0} p_i = 1.$$

b_i, c_i, v_i $(i \in D)$ are background, contents and variance of pixels observed in a subdomain $D \subseteq D_0$ of the expected distribution. The background b_i underneath a diffraction spot is often assumed to be a constant which is estimated from the neighbourhood around the reflection. Determination of reflection intensities by profile fitting has a long tradition (Diamond, 1969; Ford, 1974; Kabsch, 1988*b*; Otwinowski, 1993). Implementations of the method differ mainly in their assumptions about the variances v_i. Ford uses constant variances, which works well for films, which have a high intrinsic background. In *XDS*, which was originally designed for a multiwire detector, $v_i \propto p_i$ was assumed, which results in a straight summation of background-subtracted counts within the expected profile region, $I = \sum_{i\in D}(c_i - b_i)/\sum_{i\in D} p_i$. This particular simple formula is very satisfactory for the low background typical of these detectors. For the general case, however, better results can be obtained by using $v_i = b_i + Ip_i$ for the pixel variances as shown by Otwinowski and implemented in *DENZO* and in the later version of *XDS*. Starting with $v_i = b_i$, the intensity is now found by an iterative process which is terminated if the new intensity estimate becomes negative or does not change within a small tolerance, which is usually reached after three cycles. It can be shown that the solution thus obtained is unique.

11.3.4. Scaling

Usually, many statistically independent observations of symmetry-related reflections are recorded in the rotation images taken from one or several similar crystals of the same compound. The squared structure-factor amplitudes of equivalent reflections should be equal and the idea of scaling is to exploit this *a priori* knowledge to determine a correction factor for each observed intensity. These correction factors compensate to some extent for effects such as radiation damage, absorption, and variations in detector sensitivity and exposure times, as well as variations in size and disorder between different crystals.

The usual methods of scaling split the data into batches of roughly the same size, each covering one or more adjacent rotation images, and then determine a single scaling factor for all reflections in each batch. Neighbouring reflections may then receive quite different corrections if they are assigned to different batches. Since the selection of batch boundaries is to some extent arbitrary, a more continuous correction function would be preferable. This function could be modelled analytically (for example by using spherical harmonics) or empirically, as implemented in *XSCALE* and described below.

For each reflection, observational equations are defined as

$$\psi_{hl\alpha} = (I_{hl} - g_\alpha I_h)/\sigma_{hl}.$$

The subscript h represents the unique reflection indices and l enumerates all symmetry-related reflections to h. By definition, the unique reflection indices have the largest h, then k, then l value occuring in the set of all indices related by symmetry to the original indices, including Friedel mates. Thus, two reflections are symmetry-related if and only if their unique indices are identical. I_h is the unknown 'true' intensity and I_{hl}, σ_{hl} are symmetry-related observed intensities and their standard deviations, respectively. The subscript α denotes the coordinates at which the scaling function g_α should be evaluated. As implemented in *XDS* and *XSCALE*, $\alpha = 1, \ldots, 9$ denotes nine positions uniformly distributed in the detector plane at the beginning of data collection, $\alpha = 10, \ldots, 18$ the same positions on the detector but after the crystal has been rotated by, say, 5°, and so on. The scaling factors g_α and the estimated intensities I_h are found at the minimum of the function

$$\Psi = \sum_{hl\alpha} w_{hl\alpha}\Psi^2_{hl\alpha}.$$

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references